

COHESION IN TRANSITION METALS UNDER HIGH PRESSURES

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A bonding theory for transition metals has been derived up to a computing program. The theory permits good to very good predictions for the total energy, static pressure, Hugoniot pressure, bulk modulus and its pressure derivative as a function of the relative volume. A combination of the pseudopotential theory with the muffin tin orbital theory serves for the calculation of the cohesive properties of d-state metals. The coupling of free-electron states and local, nonoverlapping d-states is also taken into account. The nearly free s-electrons contribute four terms to the energy: kinetic energy, exchange energy, a Madelung term, and its pseudopotential correction. On this basis, two new models are developed. In both models, it is assumed that the ratio K of nearest-neighbor separation to the atomic radius is independent of the pressure p . In the first model (MB) [15], the additional condition is that a redistribution of electrons between the two outer shells, which are not completely filled, is permissible. In the second model (MC) [14], the requirement is that for elements with a full d-shell the theoretical bulk modulus at vanishing pressure is exactly equal to the experimental bulk modulus at vanishing pressure and that the pressure exactly vanishes for the zero-pressure volume. In a comparison with various kinds of measured values, MC proves to be very successful for all d-state metals except Sc, Ti and Nb. In both models, the initial pressure derivative B_1 of the bulk modulus is not an input parameter, but can be predicted. The theories are applicable in a unified manner to all transition metals.

Below, only the final results for the two equations of state will be given. For the calculation of the static pressure p as a function of the volume Ω , we have, according to model B:

$$\begin{aligned}
 p = & -c_1 \frac{5}{3} b(a\Omega^{-1} + b)^{2/3} + c_1 \frac{2}{3} (a\Omega^{-1} + b)^{5/3} + c_2 \frac{4}{3} b(a\Omega^{-1} + b)^{1/3} - \\
 & -c_2 \frac{1}{3} (a\Omega^{-1} + b)^{4/3} + 2b(a + b\Omega) (c_3\Omega^{-1/3} - c_4\Omega^{-1}) - \\
 & - (a + b\Omega)^2 (c_3 \frac{1}{3} \Omega^{-4/3} - c_4\Omega^{-2}) - \\
 & - c_5 \left[1 - \frac{c - b\Omega}{10} \right] \left(\frac{5}{3} c\Omega^{-8/3} - \frac{2}{3} b\Omega^{-5/3} \right) + \\
 & + c_5 \frac{1}{10} b(c - b\Omega)\Omega^{-5/3} + c_6 \left(\frac{8}{3} c\Omega^{-11/3} - \frac{5}{3} b\Omega^{-8/3} \right) \quad (1)
 \end{aligned}$$

with

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$$\begin{aligned}
c_1 &\equiv (2.210) \frac{\hbar^2}{2m} \left(\frac{4\pi}{3}\right)^{2/3}; & c_4 &\equiv 2\pi e^2 r_c^2; \\
c_2 &\equiv (0.916) \frac{e^2}{2} \left(\frac{4\pi}{3}\right)^{1/3}; & c_5 &\equiv (30.9)n^{1/2} \frac{\hbar^2}{2m} \frac{r_d^3}{K^5} \left(\frac{4\pi}{3}\right)^{5/3}; \\
c_3 &\equiv \frac{e^2}{2} \alpha \left(\frac{4\pi}{3}\right)^{1/3}; & c_6 &\equiv (11.40)n \frac{\hbar^2}{m} \frac{r_d^6}{K^8} \left(\frac{4\pi}{3}\right)^{8/3}. \quad (2)
\end{aligned}$$

\hbar is the Planck quantum of action, divided by 2π , e is the charge and m the mass of the electron, $\alpha = 1.8$ is fixed, r_c is the Ashcroft empty-core radius, r_d is the d-state radius of the free atom, n is the number of nearest neighbors of an atom. The three parameters a , b and c denote

$$a = Z_s^0 - Z_s^{0'} \Omega_0, \quad b = Z_s^{0'}, \quad c = Z - Z_s^0 + Z_s^{0'} \Omega_0 \quad (3)$$

where Z_s^0 is the number of s-electrons at vanishing pressure, $Z_s^{0'}$ the volume derivative of the number of s-electrons at vanishing pressure and Ω_0 the volume at zero-pressure. Higher derivatives were equated to zero. However, according to model C,

$$p = \frac{2}{3} k_1 \Omega^{-5/3} - \frac{1}{3} k_2 \Omega^{-4/3} + k_3 \Omega^{-2} - \frac{5}{3} b_n k_4 \Omega^{-8/3} + \frac{8}{3} c_n k_5 \Omega^{-11/3} \quad (4)$$

with

$$\begin{aligned}
k_1 &\equiv (2.210) \frac{\hbar^2}{2m} Z_s^{5/3} \left(\frac{4\pi}{3}\right)^{2/3}, \\
k_2 &\equiv \frac{e^2}{2} [(0.916) Z_s^{4/3} + \alpha Z_s^2] \left(\frac{4\pi}{3}\right)^{1/3}, \\
k_3 &\equiv 2\pi e^2 r_c^2 Z_s^2, \\
k_4 &\equiv Z_d \left(1 - \frac{Z_d}{10}\right) (30.9) n^{1/2} \frac{\hbar^2}{2m} \frac{r_d^3}{K^5} \left(\frac{4\pi}{3}\right)^{5/3}, \\
k_5 &\equiv Z_d n (11.40) \frac{\hbar^2}{m} \frac{r_d^6}{K^8} \left(\frac{4\pi}{3}\right)^{8/3} \quad (5)
\end{aligned}$$

where Z_d is the number of d-electrons, Z_s that of the nearly free s-electrons. If the requirement is that, with equation (4), the pressure exactly vanishes for $\Omega = \Omega_0$ and that the theoretical bulk modulus is in precise agreement with the experimental bulk modulus, you obtain $b_n \equiv b_{MA}$ and $c_n \equiv c_{MA}$. If this obvious requirement is made only for elements with full d-shells, the equation of state for model C is obtained by using equation (4) with

$$b_n = (4Z_d)^{-1} [(5Z_d - 10) b_{MA} + 10 - Z_d] \quad (6)$$

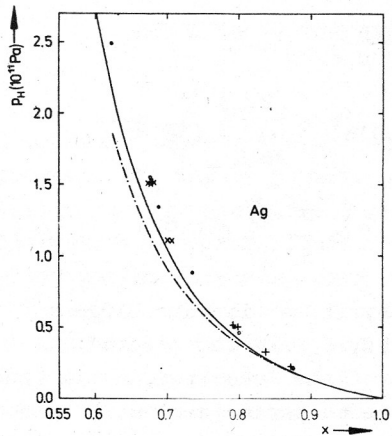


Fig. 1. Hugoniot pressure p_H versus relative volume x for silver. Theory MB is dash-dotted, theory MC continuous. The experimental values after [8] are solid circles, those after [9] upright crosses, those after [7] are oblique crosses and those after [1] empty circles.

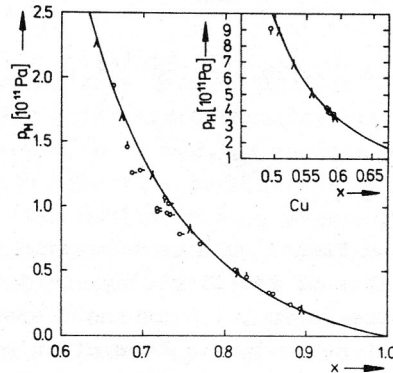


Fig. 3. Hugoniot pressure p_H versus relative volume for copper. The experimental values have been taken from [8]. The curve has been calculated according to the theory MC.

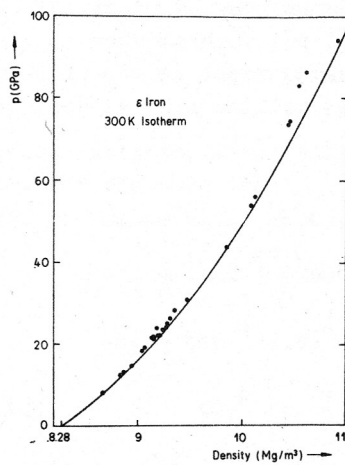


Fig. 2. Static high-pressure data p versus density for ϵ iron. The curve has been calculated according to theory MC. The solid circles represent experimental results from [5] and [6] plotted after [2].

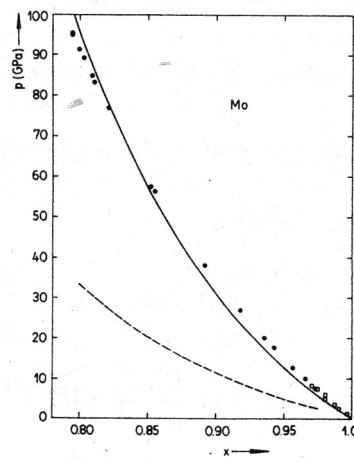


Fig. 4. Pressure p versus relative volume x . A comparison of the equations of state of MC (full curve) and of Wills and Harrison [16] (dashed curve) with experimental results for molybdenum. Full circles after [4], open squares after [3].

and

$$c_n = (4Z_d)^{-1} [(5Z_d - 10) c_{MA} + 10 - Z_d] . \quad (7)$$

The formulas have been derived in [14] and [15]. These papers are developments based on [11] and [12]. It is to be expected that the new theories [14] and [15] are better suited than [10, 13] for deriving the pressure dependence of melting temperature. In the following, a comparison with measured data is to demonstrate the applicability of the theories.

References

- [1] L.V. Altshuler, K.K. Krupnikov, and M.I. Brazhnik, Soviet Phys. - J. exper. theor. Phys. 34 (1958) 614
- [2] J.M. Brown and R.G. McQueen in: S. Akimoto, High-Pressure Research in Geophysics. Advances in Earth and Planetary Sciences 12 (1982) 611
- [3] Li Chung Ming and M.H. Manghnani, J. Appl. Phys. 49 (1978) 208
- [4] H.K. Mao, P.M. Bell, J.W. Shaner, and D.J. Steinberg, J. Appl. Phys. 49 (1978) 3276
- [5] H.K. Mao, W.A. Bassett, and T. Takahashi, J. Appl. Phys. 38 (1967) 272
- [6] H.K. Mao and P.M. Bell, J. Geophys. Res. 84 (1979) 4533
- [7] R.G. McQueen and S.P. Marsh, J. Appl. Phys. 31 (1960) 1253
- [8] M. Van Thiel, Compendium of Shock Wave Data Report, UCRL - 50108, Lawrence Livermore Laboratory, University of California 1977
- [9] J.M. Walsh, M.H. Rice, R.G. McQueen, and F.L. Yarger, Phys. Rev. 108 (1957) 169
- [10] U. Walzer, Exp. Techn. Phys. 31 (1983) 33
- [11] U. Walzer, High Temp. - High Pressures 16 (1984) 335
- [12] U. Walzer, Phys. Status Solidi B 125 (1984) 55
- [13] U. Walzer, Exp. Techn. Phys. 32 (1984) 281
- [14] U. Walzer, High Temp. - High Pressures 19 (1987) 161
- [15] U. Walzer, Phys. Status Solidi B 140 (1987) 377
- [16] J.M. Wills and W.A. Harrison, Phys. Rev. B 28 (1983) 4363